Twistors for $Lw_{1+\infty}$ symmetry in 4d gravity An open sigma model for celestial gravity

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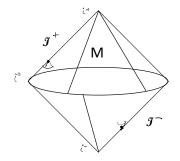
Rencontres Théoriciennes, Paris November 10, 2022

We consider 4d pure gravity in split signature. Two parts:

- Adapt Math.DG/0504582, Duke Math (2007), LeBrun & M. to present global SD gravity in split signature so as to manifest Strominger's celestial Lw_{1+∞} symmetry.

Holography from null infinity, and amplitudes

- Celestial Holography seeks to find boundary theory that constructs 4d gravity from f.
- Newman '70's: tries to rebuild space-time from 'cuts' of 𝒯.
- Yields instead 'H-space' a complex self-dual space-time.



- ▶ Penrose: \rightarrow asymptotic Twistor space $P\mathscr{T} \sim \mathbb{CP}^3$, the *nonlinear graviton*.
- Embodies integrability of SD sector.
- Chiral sigma models in twistor space give full 4d gravity S-matrix expanding around SD sector; manifests Lw_{1+∞} symmetry.



Gravity amplitudes at MHV (--+...+helicity)

Scatter *n* gravitons with momenta k_i , i = 1, ... n.

- ▶ In 2-component spinors, null momenta $k_{i\alpha\dot{\alpha}} = \kappa_{i\alpha}\kappa_{i\dot{\alpha}}$.
- Scaling of spinor $\kappa_{i\alpha}$ encodes polarization of *i*th graviton.
- Compact spinor helicity notation:

$$\langle 1\,2\rangle := \kappa_{1\alpha}\kappa_2^\alpha\,,\; [1\,2] := \kappa_{1\dot\alpha}\kappa_2^{\dot\alpha}\,,\; \langle 1|2|3] = \kappa_{1\alpha}\textit{k}_2^{\alpha\dot\alpha}\kappa_{3\dot\alpha}\,.$$

► Hodges 2012 MHV formula, defines $n \times n$ matrix:

$$\mathbb{H}_{ij} = \begin{cases} \frac{[ij]}{\langle ij \rangle} & i \neq j \\ -\sum_{k} \frac{[ik]}{\langle ik \rangle} & i = j. \end{cases}$$

► Then:

$$\mathcal{M}(1,\ldots,n) = \langle 12 \rangle^6 \det{}' \mathbb{H} \,\, \delta^4(\sum_i k_i)$$
 Why???
$$\mathcal{M} = \langle V_1 \ldots V_{n-2} \rangle.$$

Flat holography: the split signature story from \mathscr{I}

A celestial torus

Now $\mathscr{I} = \mathbb{R} \times S^1 \times S^1$ with real coords $(u, \lambda, \tilde{\lambda})$, $\lambda = \lambda_1/\lambda_0$.

$$\label{eq:ds2} \textit{ds}^2 = \frac{1}{R^2} \left(\textit{dudR} - \textit{d}\lambda \textit{d}\tilde{\lambda} + \textit{R}\sigma \textit{d}\tilde{\lambda}^2 + \textit{R}\tilde{\sigma} \textit{d}\lambda^2 + \ldots \right) \,,$$

where R = 1/r, and $\mathscr{I} = \{R = 0\}$.

- ▶ The σ , $\tilde{\sigma}$ are now *real* asymptotic *shears* that encode gravitational data.
- ightharpoonup σ encodes self-dual (SD) sector and $\tilde{\sigma}$ the ASD sector.
- ▶ Split signature \sim real 'twistors' = totally null ASD 2-planes.
- ▶ Twistors intersect \mathscr{I} in null geodesic circles in $\lambda = \text{const.}$ planes:

$$u = Z(\lambda, \tilde{\lambda}), \qquad \frac{\partial^2 Z}{\partial \tilde{\lambda}^2} = \sigma(Z, \lambda, \tilde{\lambda}).$$

▶ We will show how twistor construction encodes $(\sigma, \tilde{\sigma})$ into twistor data $h(U), \tilde{h}(\tilde{U})$ encoding $Lw_{1+\infty}$ action.

SD sector arises by solving open disk chiral sigma model, and gives formulae for perturbations about SD sector.

Conformal self-duality in 4d, split signature

Recall on 4d manifold (M^4, g) ,

$$\Omega_{\textit{M}}^2 = \begin{pmatrix} \Omega^{2+} \\ \oplus \\ \Omega^{2-} \end{pmatrix} \,, \qquad \text{Riem} = \begin{pmatrix} \text{Weyl}^+ + \textit{S}\delta & \text{Ricci}_0 \\ \text{Ricci}_0 & \text{Weyl}^- + \textit{S}\delta \end{pmatrix}.$$

This talk: focus on Ricci = $0 = \text{Weyl}^-$, so Ω^{2-} is flat.

Conformal group = SO(3,3) acts on global models:

▶ Conformally flat models: $S^2 \times S^2$ or $S^2 \times S^2/\mathbb{Z}_2$:

$$\label{eq:ds2} \textit{ds}^2 = \Omega^2 (\textit{ds}^2_{\textit{S}^2_{\textbf{x}}} - \textit{ds}^2_{\textit{S}^2_{\textbf{y}}}) \,,$$

Coordinates $(\mathbf{x}, \mathbf{y}) \in \mathbb{R}^3 \times \mathbb{R}^3$, $|\mathbf{x}| = |\mathbf{y}| = 1$.

- $ightharpoonup \mathbb{Z}_2$ acts by $(\mathbf{x},\mathbf{y}) \to (-\mathbf{x},-\mathbf{y})$.
- For flat $\Lambda = 0 : \Omega \sim \frac{1}{x_2 v_2}$, and

$$\mathscr{I} = \{x_3 = y_3\} = \mathbb{R} \times S^1 \times S^1.$$

(For
$$\Lambda \neq 0$$
: $\Omega \sim 1/y_3$, and $\mathscr{I} = S^2 \times S^1$.)

α and β -surfaces and the Zollfrei condition

The split signature conformally flat metric

$$ds^2 = \Omega^2 (ds_{S_x^2}^2 - ds_{S_y^2}^2),$$

admits a 3-parameter family of β -planes denoted by $\mathbb{PT}_{\mathbb{R}}$:

respectively totally null ASD S2s given by

$$\mathbf{x} = A\mathbf{y}$$
, $A \in SO(3) = \mathbb{RP}^3$.

- ▶ Weyl⁻ = $0 \Rightarrow \beta$ -planes survive as β -surfaces.
- \triangleright β -surfaces are projectively flat.
- ▶ If compact, β -surfaces are necessarily S^2 or \mathbb{RP}^2 .
- ▶ Null geodesics are projectively \mathbb{RP}^1 s or double cover.

Following Guillemin we define:

Definition

An indefinite space (M^d, g) is (strongly) Zollfrei if all null geodesics are embedded S^1 s (of same projective length).



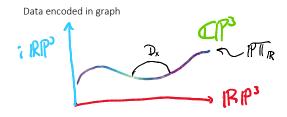
Conformally self-dual case

Theorem (LeBrun & M. [Duke Math J. 2007, math.dg/0504582.)

Let $(M^4, [g])$ be Zollfrei with SD Weyl-curvature. Then either

- $M = S^2 \times S^2/\mathbb{Z}_2$ with the standard conformally flat conformal structure, or
- ▶ $M = S^2 \times S^2$ and there is a 1 : 1-correspondence between
 - 1. SD conformal structures on $S^2 \times S^2$ near flat model &
 - 2. Deformations $\mathbb{PT}_{\mathbb{R}}$ of standard embedding of $\mathbb{RP}^3 \subset \mathbb{CP}^3$ modulo reparametrizations of \mathbb{RP}^3 and $PGL(4,\mathbb{C})$ on \mathbb{CP}^3 .

The space $\mathbb{PT}_{\mathbb{R}} = \{\beta \text{ surfaces in } M\} = \text{graph of } F : \mathbb{RP}^3 \to \mathbb{R}^3 \text{ in some neighbourhood } U \simeq \mathbb{R}^3 \times \mathbb{RP}^3 \subset \mathbb{CP}^3 \text{ of } \mathbb{RP}^3$:



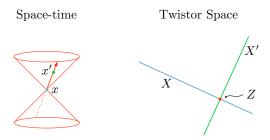
Reconstruction of M from twistor space $\mathbb{PT}_{\mathbb{R}}$

Each $x \in M \leftrightarrow$ holomorphic disc $\mathbb{D}_x \subset \mathbb{CP}^3$ with $\partial \mathbb{D}_x \subset \mathbb{PT}_{\mathbb{R}}$.

- ▶ \mathbb{D}_X generates the degree-1 class in $H_2(\mathbb{CP}^3, \mathbb{PT}_{\mathbb{R}}, \mathbb{Z}) = \mathbb{Z}$.
- ▶ Reconstruct M from $\mathbb{PT}_{\mathbb{R}}$ space of all such disks:

 $M = \{ \text{Moduli of degree-1 hol. disks: } \mathbb{D}_x \subset \mathbb{CP}^3, \partial \mathbb{D}_x \subset \mathbb{PT}_{\mathbb{R}} \}$

- Gives compact 4d moduli space
- ▶ *M* admits a conformal structure for which $\partial \mathbb{D}_x \cap \partial \mathbb{D}_{x'} = Z$ means that x, x' sit on same β -plane:



Restriction to Einstein vacuum case

Adapting Penrose nonlinear graviton (1976) to split signature

Which $\mathbb{PT}_{\mathbb{R}} \subset \mathbb{CP}^3$ give SD Einstein $g \in [g]$ on $S^2 \times S^2$?

- Let $Z^A = (\lambda_{\alpha}, \mu^{\dot{\alpha}})$, $\alpha = 0, 1, \dot{\alpha} = \dot{0}, \dot{1}$ be homogenous coordinates for \mathbb{CP}^3 .
- Introduce Poisson structure and 1-form

$$\{f, g\} := arepsilon^{\dot{lpha}\dot{eta}} rac{\partial f}{\partial \mu^{\dot{lpha}}} rac{\partial g}{\partial \mu^{\dot{eta}}} = \left[rac{\partial f}{\partial \mu} rac{\partial g}{\partial \mu}
ight] \,, \ heta := \epsilon^{lphaeta} \lambda_{lpha} d\lambda_{eta} = \langle \lambda d\lambda
angle$$

of rank 2 and homogeneity degree -2 and 2 respectively.

Theorem

A vacuum $g \in [g]$ exists when $\theta|_{\mathbb{PT}_{\mathbb{P}}} \& \{,\}_{\mathbb{PT}_{\mathbb{P}}}$ are real.

Poisson diffeos of plane & $Lw_{1+\infty}$

 $W_N =$ higher spin symmetries in 2d CFT [Zamolodchikov 1980s]. For $N \to \infty$ classical limit $w_\infty =$ Poisson diffeos of plane [Hoppe].

▶ Plane has coords $\mu^{\dot{\alpha}}$, $\dot{\alpha} = \dot{0}$, $\dot{1}$ with Poisson bracket

$$\{f,g\} := arepsilon^{\dot{lpha}\dot{eta}} rac{\partial f}{\partial \mu^{\dot{lpha}}} rac{\partial g}{\partial \mu^{\dot{lpha}}} \,, \qquad arepsilon^{\dot{lpha}\dot{eta}} = arepsilon^{[\dot{lpha}\dot{eta}]}, \quad arepsilon^{\dot{0}\dot{1}} = 1 \,.$$

▶ Basis of $w_{1+\infty} \leftrightarrow \text{polynomial hamiltonians}$

$$w_m^p = (\mu^{\dot{0}})^{p-m-1} (\mu^{\dot{1}})^{p+m-1}, \qquad |m| \le p-1, \quad 2p-2 \in \mathbb{N}$$

▶ Poisson brackets \leftrightarrow commutation relations of $w_{1+\infty}$:

$$\{w_m^p, w_n^q\} = (2(p-1)n - 2(q-1)m)w_{m+n}^{p+q-2}$$
.

▶ Loop algebra $Lw_{1+\infty}$, loop coord $\frac{\lambda_1}{\lambda_0} = \tan \frac{\theta}{2}$, generators

$$g_{m,r}^p = w_m^p e^{ir\theta}, \qquad r \in \mathbb{Z}.$$

Poisson brackets now

$$\{g_{m,r}^p, g_{n,s}^q\} = (2(p-1)n - 2(q-1)m)g_{m+n,r+s}^{p+q-2}$$

This is the structure preserving diffeomorphism group of $\mathbb{PT}_{\mathbb{R}}$.

Generating functions for Einstein embeddings

Explicitly in homogeneous coordinates:

- ▶ Let $Z^A = U^A + iV^A$, with U^A , $V^A \in \mathbb{R}^4$.
- ▶ Let h(U) be an arbtrary function of homogeneity degree 2,

$$U \cdot \frac{\partial h}{\partial U} = 2h.$$

Proposition

All 'small' Einstein vacuum twistor data $\leftrightarrow h(U)$ by setting

$$\mathbb{T}_{\mathbb{R}} = \left\{ V^{A} = \left\{ h, U^{A} \right\} \right\} = \left\{ v_{\alpha} = 0, v^{\dot{\alpha}} = \varepsilon^{\dot{\alpha}\dot{\beta}} \frac{\partial h}{\partial u^{\dot{\beta}}} \right\}$$

projectivising gives $\mathbb{PT}_{\mathbb{R}}$.

The corresponding self-dual (2,2) vacuum metrics are Zollfrei on $S^2 \times S^2$ with null \mathscr{I} modelled by $x_3 = y_3$.

The Poisson bracket underpins Strominger's $Lw_{1+\infty}$ structure, [Adamo, M., Sharma, 2110.06066.]. Here $Lw_{1+\infty}$ acts canonically on

 $\{SD \text{ gravity phase space}\} = Lw_{1+\infty}^{\mathbb{C}}/Lw_{1+\infty} \ni h(U)$

Holography: SD vacuum spaces from *I*

Twistor space can be constructed from σ at \mathscr{I} :

• At fixed λ_{α} , real twistor coords $\mu^{\dot{\alpha}}$ parametrize null geodesics $u = Z(\tilde{\lambda})$ in \mathscr{I} where

$$\partial_{\tilde{\lambda}}^2 Z = \sigma(Z, \tilde{\lambda}, \lambda).$$

Defines Zollfrei projective structure on each $\lambda = \text{const.}$.

- Flat $\sigma = 0$ case has $Z = \mu^{\dot{\alpha}} \tilde{\lambda}_{\dot{\alpha}}$.
- ► In general ∃ nonlinear correspondence [Lebrun & M, JDiffGeom. '02]:

{Zollfrei proj. str.
$$\leftrightarrow \sigma$$
} $\stackrel{\text{1:1}}{\longleftrightarrow}$ { $h(U)$ },

and gives $\mathscr{I} \leftrightarrow \mathbb{PT}_{\mathbb{R}} \subset \mathbb{PT}$ at each fixed λ .

In linear theory map is analogue of radon transform

$$\sigma(u, \tilde{\lambda}, \lambda) = \partial_u^2 \int_{-\infty}^{\infty} dt \ h(\mu^{\dot{\alpha}} + t\tilde{\lambda}^{\dot{\alpha}}, \lambda_{\alpha}).$$

in α -planes at $\mathscr I$ (cf inverse light-ray transform).



Examples:

- Let $T^{\alpha\dot{\alpha}}$, $T^2 = 2$, be symmetry.
- ▶ Use $T^{\alpha\dot{\alpha}}$ to eliminate dotted indices.
- ▶ So $Z^A = (\lambda_\alpha, \mu^\alpha)$, and $\{f, g\} = \varepsilon^{\alpha\beta} \frac{\partial f}{\partial \mu^\alpha} \frac{\partial g}{\partial \mu^\beta}$,
- ▶ Set $\mu^{\dot{\alpha}} = u^{\alpha} + iv^{\alpha}$, $h = h(u^{\alpha}\lambda_{\alpha}, \lambda_{\alpha})$ then define

$$\mathbb{PT}_{\mathbb{R}} = \{ \Im \lambda_{\alpha} = \mathbf{0}, \mathbf{v}^{\alpha} = \lambda^{\alpha} \dot{\mathbf{h}} \}, \qquad \dot{\mathbf{h}}(\mathbf{w}, \lambda) := \partial_{\mathbf{w}} \mathbf{h}(\mathbf{w}, \lambda).$$

For hol. disks: use λ_{α} as hgs coords & express as graphs:

$$\mu^{\alpha} = \mathbf{x}^{\alpha\beta}\lambda_{\beta} + (t + \mathbf{g}(\mathbf{x}, \lambda))\lambda^{\alpha}, \qquad \mathbf{x}^{\alpha\beta} = \mathbf{x}^{(\alpha\beta)}.$$

where

$$g(x^{\alpha\beta},\lambda) = \oint \frac{\lambda_0}{\lambda_\alpha'} \frac{1}{\langle \lambda \lambda' \rangle} \dot{h}((x^{\alpha\beta}\lambda_\alpha'\lambda_\beta',\lambda_\alpha')D\lambda'$$

Gives split signature version of Gibbons-Hawking metrics

$$ds^2 = V d\mathbf{x} \cdot d\mathbf{x} + V^{-1} (dt + \omega)^2$$
, $dV = ^* d\omega$, $V = \oint \ddot{h} D\lambda$.

$$\square_{2+1} V = 0$$
. E.g. $V = 1 + 2m/r$ for SD Schwarzschild.



Amplitudes from open chiral twistor sigma models

Represent holomorphic disks $\mathbb{D}\subset\mathbb{PT}$ with boundary $\partial\mathbb{D}\subset\mathbb{PT}_{\mathbb{R}}$ in homogeneous coordinates by

$$Z^{A}(\sigma): \mathbb{D} \to \mathbb{T}, \qquad Z^{A}|_{\sigma = \bar{\sigma}} \in \mathbb{T}_{\mathbb{R}}.$$

using disk as upper-half-plane $\mathbb{D} = \{ \sigma \in \mathbb{C}, \Im \sigma \geq 0 \}.$

▶ For k points $\sigma_i \in \mathbb{R}$, and $Z_i^A \in \mathbb{T}_{\mathbb{R}}$, $\exists ! \text{ deg } k-1 \text{ disk thru } Z_i$:

$$Z^{A}(\sigma) = \sum_{i=1}^{K} \frac{Z_{i}^{A}}{\sigma - \sigma_{i}} + M(\sigma), \qquad M(\sigma) \text{ holomorphic on } \mathbb{D}.$$

- ▶ For $Z = (\lambda_{\alpha}, \mu^{\dot{\alpha}}) \in \mathbb{T}_{\mathbb{R}}$ implies λ_{α} real.
- ► Therefore $M^A = (0, m^{\dot{\alpha}})$, but $m^{\dot{\alpha}} \neq 0$ unless h = 0.
- Action for holomorphy and boundary conditions:

$$S_D[Z(\sigma), Z_i] = \int_{\mathbb{D}} [m\,\bar{\partial}m] d\sigma + \oint_{\partial\mathbb{D}} h(Z) d\sigma$$

using *spinor-helicity* notation $[\mu \nu] := \mu_{\dot{\alpha}} \nu^{\dot{\alpha}}, \langle 1 \, 2 \rangle := \kappa_{1\alpha} \kappa_2^{\alpha}.$

Sigma model and gravity S-matrix on SD background

Amplitudes are functionals $\mathcal{M}[h, \tilde{h}_i]$ of gravitational data:

- ▶ $h \in C^{\infty}(\mathbb{PT}_{\mathbb{R}}, \mathcal{O}(2))$ for fully nonlinear SD part,
- $\tilde{h}_i \in C^{\infty}(\mathbb{PT}_{\mathbb{R}}, \mathcal{O}(-6)), i = 1, \dots, k$, ASD perturbations.
- ▶ For eigenstates of momentum $k_{i\alpha\dot{\alpha}} = \kappa_{i\alpha}\tilde{\kappa}_{i\dot{\alpha}}$ take:

$$h_{i} = \int \frac{dt}{t^{3}} \delta^{2}(t\lambda_{\alpha} - \kappa_{i\alpha}) e^{it[\mu, \tilde{\kappa}_{i}]}, \quad \tilde{h}_{i} = \int \frac{dt}{t^{-5}} \delta^{2}(t\lambda_{\alpha} - \kappa_{i\alpha}) e^{it[\mu, \tilde{\kappa}_{i}]}$$

Proposition (Adapted from [Adamo, M. & Sharma, 2103.16984] to split signature.)

The amplitude for k ASD perturbations on SD background h is

$$\mathcal{M}(h, \tilde{h}_i) = \int_{(S^1 imes \mathbb{PT}_{\mathbb{R}})^k} S^{os}_D[h, Z_i, \sigma_i] \det{}' \tilde{\mathbb{H}} \prod_{i=1}^k \tilde{h}_i(Z_i) D^3 Z_i d\sigma_i \,.$$

Here $S_D^{os}[h, Z_i, \sigma_i]$ is the on-shell Sigma model action and

$$ilde{\mathbb{H}}_{ij}(Z_i) = egin{cases} rac{\langle \lambda_i \lambda_j
angle}{\sigma_i - \sigma_j} & i
eq j \\ -\sum_{l} rac{\langle \lambda_i \lambda_l
angle}{\sigma_i - \sigma_j} \,, & i = j \,. \end{cases}$$

Ideas in proof: the complete tree-level S-matrix

- Expand $h = h_{k+1} + \ldots + h_n$ to 1st order in momentum e-states h_i to give flat background perturbative amplitude.
- On shell action expands as tree correlator

$$S_D^{os}[h_{k+1} + \ldots + h_n, Z_i, \sigma_i] = \langle V_{h_{k+1}} \ldots V_{h_n} \rangle_{tree} + O(h_i^2).$$

- ► Here the 'vertex operators' are $V_{h_i} = \int_{\partial D} h_i(\sigma_i) d\sigma_i$.
- ▶ Propagators for S_D give Poisson bracket { , }

$$\langle h_i h_j \rangle_{tree} = \frac{[\partial_\mu h_i \, \partial_\mu h_j]}{\sigma_i - \sigma_i} = \frac{[ij]}{\sigma_i - \sigma_i} h_i h_j \,, \qquad i \neq j \,.$$

Matrix-tree theorem then gives

$$\langle h_{k+1} \dots h_n \rangle_{\mathrm{tree}} = \det' \mathbb{H} \prod^n h_i, \qquad \mathbb{H}_{ij} = \frac{[ij]}{\sigma_i - \sigma_i}, \quad i \neq j \text{ etc.}$$

This is now equivalent to the Cachazo-Skinner formula.



Relation to Einstein-Hilbert action at k=2

[Adamo, M, Sharma, 2103.1239]

At k=2, $\det{}'\tilde{\mathbb{H}}$ and Mobius symmetry trivialises σ integrals so

$$\mathcal{M}[h, \tilde{h}_1, \tilde{h}_2] = \int d^2\mu_1 d^2\mu_2 \ \mathrm{e}^{i[\mu_1 \ 1] + i[\mu_2 \ 2]} \mathcal{S}^{os}_D[h, Z_1, Z_2]$$

• Writing $x^{\alpha\dot{\alpha}}=(\mu_1^{\dot{\alpha}},\mu_2^{\dot{\alpha}})$ this a space-time integral

$$\mathcal{M}[h, \tilde{h}_1, \tilde{h}_2] = \int d^4x \, e^{ik_1 \cdot x + ik_2 \cdot x} S_D^{os}[h, \mu_1, \mu_2]$$

Proposition

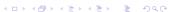
Let $\Omega(x) := S_{\mathcal{D}}^{os}[h, \mu_1, \mu_2]$. Then Ω is the Plebanskis first potential (Kahler scalar) for the SD background metric

$$ds^2 = rac{\partial^2 \Omega}{\partial \mu_1^{\dot{lpha}} \partial \mu_2^{\dot{eta}}} d\mu_1^{\dot{lpha}} d\mu_2^{\dot{eta}} \, .$$

The second variation of the Einstein-Hilbert action

$$\delta^2 S_{\text{EH}}[h, \tilde{h}_1, \tilde{h}_2] = \int d^4 x e^{i(k_1 + k_2) \cdot x} \Omega(x) = \mathcal{M}[h, \tilde{h}_1, \tilde{h}_2]$$

(Follows from Plebanski gravity action.)



Conclusions and open problems

- ▶ We have rigidity of conformally-flat SD split signature vacuum metrics with $\mathscr{I} = S^1 \times S^1 \times \mathbb{R}/\mathbb{Z}_2$.
- ▶ Have construction for split signature SD vacuum metrics on $S^2 \times S^2$ with $\mathscr{I} \simeq S^1 \times S^1 \times \mathbb{R}$ depending on smooth sections h of $\mathcal{O}(2)$ over \mathbb{RP}^3 defining deformed real slice.
- Similar results follow for $\Lambda \neq 0$ where $h \leftrightarrow 2 + 1$ signature conformal structure of $\mathscr{I} = S^2 \times S^1$.
- Reconstruction via open holomorphic discs leads to chiral open sigma model that computes gravity amplitudes.
- ► MHV formula gives theory underlying tree formalism of Bern et. al. from 1998.
- Framework gives $Lw_{1+\infty}$ action on *full amplitude*. **Slogan:** SD gravity phase space = $Lw_{1+\infty}^{\mathbb{C}}/Lw_{1+\infty}$
- Split signature twistors avoid 'lightray transform' or Čech-Dolbeult manifesting Lw_{1+∞} directly.



Thank you!