

# **Symmetries of gravitational scattering in absence of peeling**

**Based on 2407.07978 with M. Geiller and A. Laddha**

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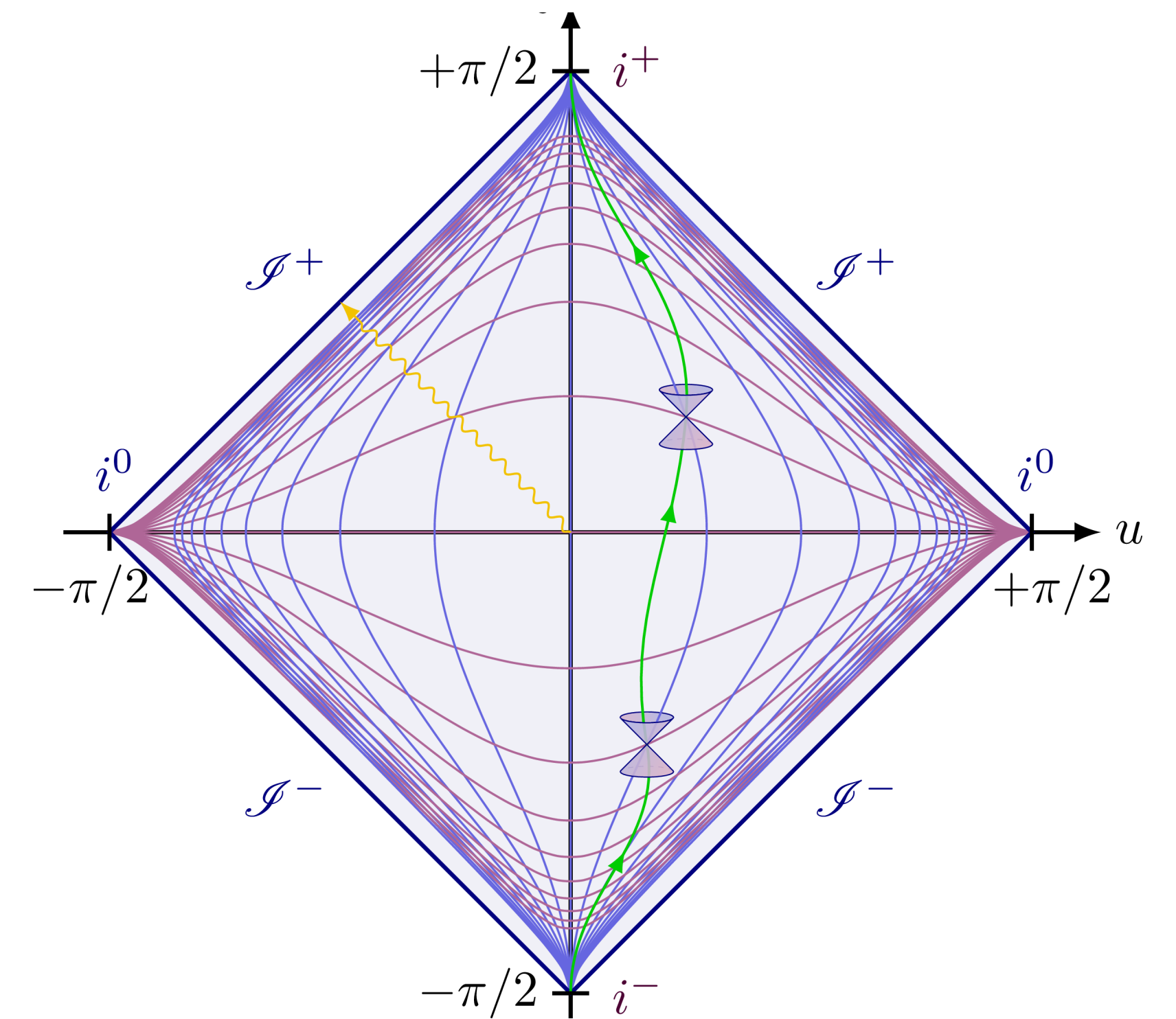
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# Introduction to Bondi gauge

# Introduction

- Asymptotic states (scattering amplitudes)
- Conformal compactification: brings the infinite boundary to a finite distance by a conformal transformation
- Penrose diagram
- Focus on radiation and therefor on future null infinity  
-> adapted gauge: Bondi gauge
- Asymptotic boundaries allow to define a notion of energy (or other observables) for gravity



from I. Neutelings on TikZ.net

# Bondi gauge [H. Bondi, M. G. J. van der Burg and A. W. K. Metzner '62; Sachs '61]

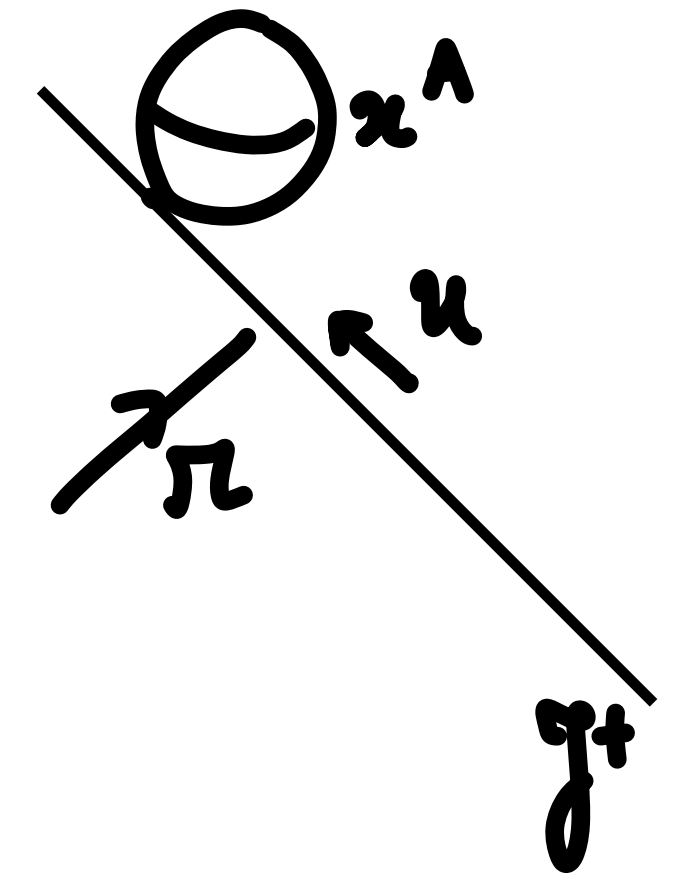
## Gauge adapted to future null infinity

1. Retarded time  $u$  is a null coordinate  $g^{uu} = 0$
2. Angular coordinates  $x^A = (\theta, \phi)$  are transverse to the null direction  $g^{uA} = 0$

$$ds^2 = e^{2\beta} \frac{V}{r} du^2 - 2e^{2\beta} du dr + \gamma_{AB} (dx^A - U^A du)(dx^B - U^B du)$$

3. Bondi determinant condition  $\partial_r \left( \frac{\sqrt{\gamma}}{r^2} \right) = 0$  and  $\gamma_{AB} = \mathcal{O}(r^2)$

Solving the eom will determine the radial dependence of  $(V, U^A, \beta)$



$$ds^2 = e^{2\beta} \frac{V}{r} du^2 - 2e^{2\beta} du dr + \gamma_{AB} (dx^A - U^A du)(dx^B - U^B du)$$

Asymptotically (assuming  $\beta_0 = 0 = U_0^A = \partial_u q_{AB}$ ),

$$\gamma_{AB} = r^2 q_{AB} + r C_{AB} + o(r)$$

$C_{AB}$  is the shear,  $M$  the mass,  $R$  the Ricci of  $q_{AB}$

$$\beta = -\frac{1}{r^2} C_{AB} C^{AB} + \mathcal{O}(r^3)$$

$$U^A = -\frac{1}{2r^2} D_B C^{AB} + o(r^{-3})$$

$$V = -\frac{R}{2} + 2M + o(1)$$

# BMS & Asymptotic symmetries

- Asymptotic symmetries preserved the choice of gauge and boundary conditions
- The generator associated to the asymptotic symmetries are **charges**, which carries physical information such as the mass of spacetimes, ...
- They satisfy an algebra which is an entry of the holographic dictionary -> use in bottom-up approach of holography

# BMS & Asymptotic symmetries

$$\{g_{BC-Bondi} + \mathcal{L}_\xi g_{BC-Bondi}\} = \{g_{BC-Bondi}\}$$

$$\text{Ex: } g_{rr} = 0 \Rightarrow \mathcal{L}_\xi g_{rr} = 0 \Rightarrow \xi^u = f(u, x^A)$$

Imposing the rest of gauge conditions and boundary conditions we obtain

$$\xi^u = f(u, x^A) = T(x^A) + u W(x^A)$$

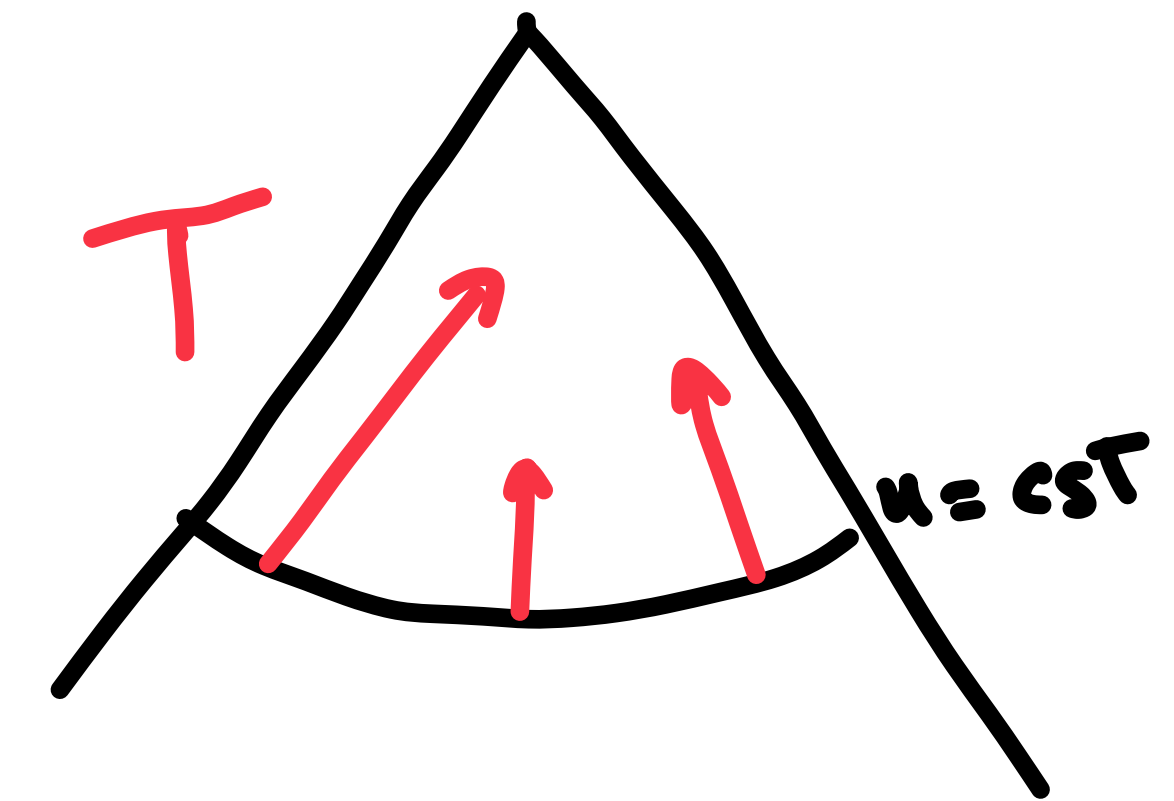
$$\xi^r = -r W(x^A) + \frac{r}{2}(U^A \partial_A f - D_A \xi^A) = -r W(x^A) + \frac{1}{2} \Delta f + \dots$$

$$\xi^A = Y^A - \frac{1}{r} \partial^A f + \dots$$

$T$  = supertranslation

$Y^A$  = superrotations

$W$  = Weyl rescaling





# BMS algebra

$$\xi(T(x^A), W(x^A), Y^A(x^A))$$

$$T_{12} = T_1 W_2 + Y_1^A \partial_A T_2 - (1 \leftrightarrow 2)$$

$$[\xi(T_1, W_1, Y_1^A), \xi(T_2, W_2, Y_2^A)] = \xi(T_{12}, W_{12}, Y_{12}^A) \quad \text{with} \quad W_{12} = Y_1^A \partial_A W_2 - (1 \leftrightarrow 2)$$

$$Y_{12}^A = Y_1^B \partial_B Y_2^A - (1 \leftrightarrow 2)$$

( $\text{Diff}(S^2)(Y^A)$  semi-direct sum  $\mathbb{R}(W)$ ) semi-direct sum  $\mathbb{R}(T)$  = BMSW

[Barnich-Troessaert '10; Freidel, Oliveri Pranzetti, Speciale '21]

If  $\delta_\xi \sqrt{q} = 0$ , then  $W = \frac{1}{2} D_A Y^A$  and we talked of the generalized BMS group

[Campiglia-Laddha '15, Compère, Fiorucci, Ruzziconi '18]

If  $\delta_\xi q_{AB} = 0$  and  $Y^A$  globally well-defined, you land on global BMS group

[H. Bondi, M. G. J. van der Burg and A. W. K. Metzner '62; Sachs '61]

# Peeling property

# Newman-Penrose formalism

- Introduction of a null tetrad  $\ell^\mu = \partial_r, n^\mu = \partial_u - \frac{R}{4}\partial_r + \mathcal{O}(r^{-1}),$   
 $m^\mu = \frac{1}{r}m_0^A\partial_A + \mathcal{O}(r^{-2}), \bar{m}^\mu = \frac{1}{r}\bar{m}_0^A\partial_A + \mathcal{O}(r^{-2})$
- $\sigma = -\frac{1}{2}C_{AB}m_0^Am_0^B$
- contraction of the Weyl tensor with the null tetrad = Weyl scalars = 5 complex numbers  $\Psi_n$

# Weyl Scalars

$\Psi_4, \Psi_3$  carries information about the radiation ( $\Psi_4^0 \propto \partial_u^2 \sigma$ ,  $\Psi_3^0 \propto m_0^A \partial_A \partial_u \sigma$ )

$\text{Re}(\Psi_2^0)$  the mass,

$\Psi_1^0$  the angular momentum

$\Psi_0$  the higher spin charges

**Peeling property**  $\Psi_n = \frac{1}{r^{5-n}} \Psi_n^0 + \mathcal{O}(r^{n-6})$

# BC compatible with the peeling = Asympt. Simple

$$\gamma_{AB} = r^2 q_{AB} \sqrt{1 + \frac{\mathcal{C}_{AB} \mathcal{C}^{AB}}{2r^2}} + r \mathcal{C}_{AB}$$

$$r \mathcal{C}_{AB} = r C_{AB} + \mathbf{0} + \sum_{n=1}^{\infty} \frac{1}{r^n} E_{AB}^n$$

$$U^A = -\frac{1}{2r^2} D_B C^{AB} + \frac{1}{r^3} N^A + \mathcal{O}(r^{-4})$$

$$V = -\frac{R}{2} + 2M + \mathcal{O}(r^{-1})$$

$$\Psi_4 = \frac{1}{r} \frac{1}{2} \partial_u^2 C_{AB} \bar{m}_0^A \bar{m}_0^B + \mathcal{O}(r^{-2})$$

$$\Psi_3 = \frac{1}{r^2} \frac{1}{2} D^B (\partial_u C_{AB}) \bar{m}_0^A + \mathcal{O}(r^{-3})$$

$$\Psi_2 = \frac{1}{r^3} \left( M + \frac{1}{16} \partial_u C_{AB} C^{AB} + i\tilde{M} \right) + \mathcal{O}(r^{-4})$$

$$\Psi_1 = \frac{1}{r^4} \left( -\frac{3}{2} N_A + \frac{3}{32} D_A C_{BC} C^{BC} + \frac{3}{4} C_{AB} D_C C^{BC} \right) m_0^A + \mathcal{O}(r^{-5})$$

$$\Psi_0 = \frac{1}{r^5} \left( 3E_{AB}^1 - \frac{3}{16} \partial_A C_{BC} C^{BC} \right) m_0^A m_0^B + \mathcal{O}(r^{-6})$$

$$\Psi_n = \frac{1}{r^{5-n}} \Psi_n^0 + \mathcal{O}(r^{n-6})$$

**Absence of the peeling**

# Logarithmically Asympt. Flat

$$\gamma_{AB} = r^2 q_{AB} \sqrt{1 + \frac{\mathcal{C}_{AB} \mathcal{C}^{AB}}{2r^2}} + r \mathcal{C}_{AB}$$

$$r \mathcal{C}_{AB} = r C_{AB} + D_{AB} + \sum_{n=1}^{\infty} \sum_m^{n+1} \frac{\ln(r)^m}{r^n} E_{AB}^{n,m}$$

$$U^A = -\frac{1}{2r^2} D_B C^{AB} + \frac{1}{r^3} (N^A + \dots \ln r) + o(r^{-3})$$

$$V = -\frac{R}{2} + 2M + o(r^0)$$

$$\partial_u D_{AB} = 0$$

$$\Psi_4 = \frac{1}{r} \Psi_4^0 + \mathcal{O}(r^{-2})$$

$$\Psi_3 = \frac{1}{r^2} \Psi_3^0 + \mathcal{O}(r^{-3})$$

$$\Psi_2 = \frac{1}{r^3} \Psi_2^0 + \mathcal{O}(r^{-4})$$

$$\Psi_1 = \frac{1}{r^4} (\Psi_1^0 + D^B D_{AB} m_0^B \ln r) + o(r^{-4})$$

$$\Psi_0 = \frac{1}{r^4} D_{AB} m_0^A m_0^B$$

$$+ \frac{1}{r^5} \left( (\Psi_0^0 + \dots) + (3E^{1,1} + \dots) \ln r + 3E_{AB}^{1,2} \ln r^2 \right)$$

$$+ o(r^{-5})$$

[Winicour '85; P. T. Chrusciel, M. A. H. MacCallum and D. B. Singleton '93; Kroon '98]

[L. Kehrberger and H. Masaoood '24]

# Why?

## General relativity

- There has been a debate on “Peeling or not peeling” [Friedrich '17]
- Blanchet’s theorem: past stationarity & no incoming radiation -> No violation of peeling at future null infinity
- **BUT** typical radiative fields sourced by matter stress tensor yields violation of peeling (no past stationarity).  
For instance: hyperbolic encounters [Damour '86, Christodoulou '02, Kehrberger '21->'24, ...]

*What is the impact of violation of peeling on the symmetries, the charges, fluxes?*



# Why?

## Soft theorems

- In a scattering process, the gravitational field has remarkable universal properties at the late/early time.
- Displacement memory effect: gravitational field is dominated to by a static mode which is fixed only by incoming and out coming momenta of the scattered particles at late/early time

This is the same thing as Weinberg soft theorem.

$$C_{AB} = C_{AB}^0(x^A) + \frac{1}{u} C_{AB}^1(x^A) + o(u^{-1})$$

- Saha, Sahoo, Sen shows that the shear  $C_{AB}^1$  is also universal (incoming and out coming momenta of the scattered particles).

This is the same thing as logarithmic soft theorem

# Why?

## Christodoulou's heuristic argument

Using the eom, Christodoulou showed that it will precisely yield a violation of peeling at the order  $\Psi_1$

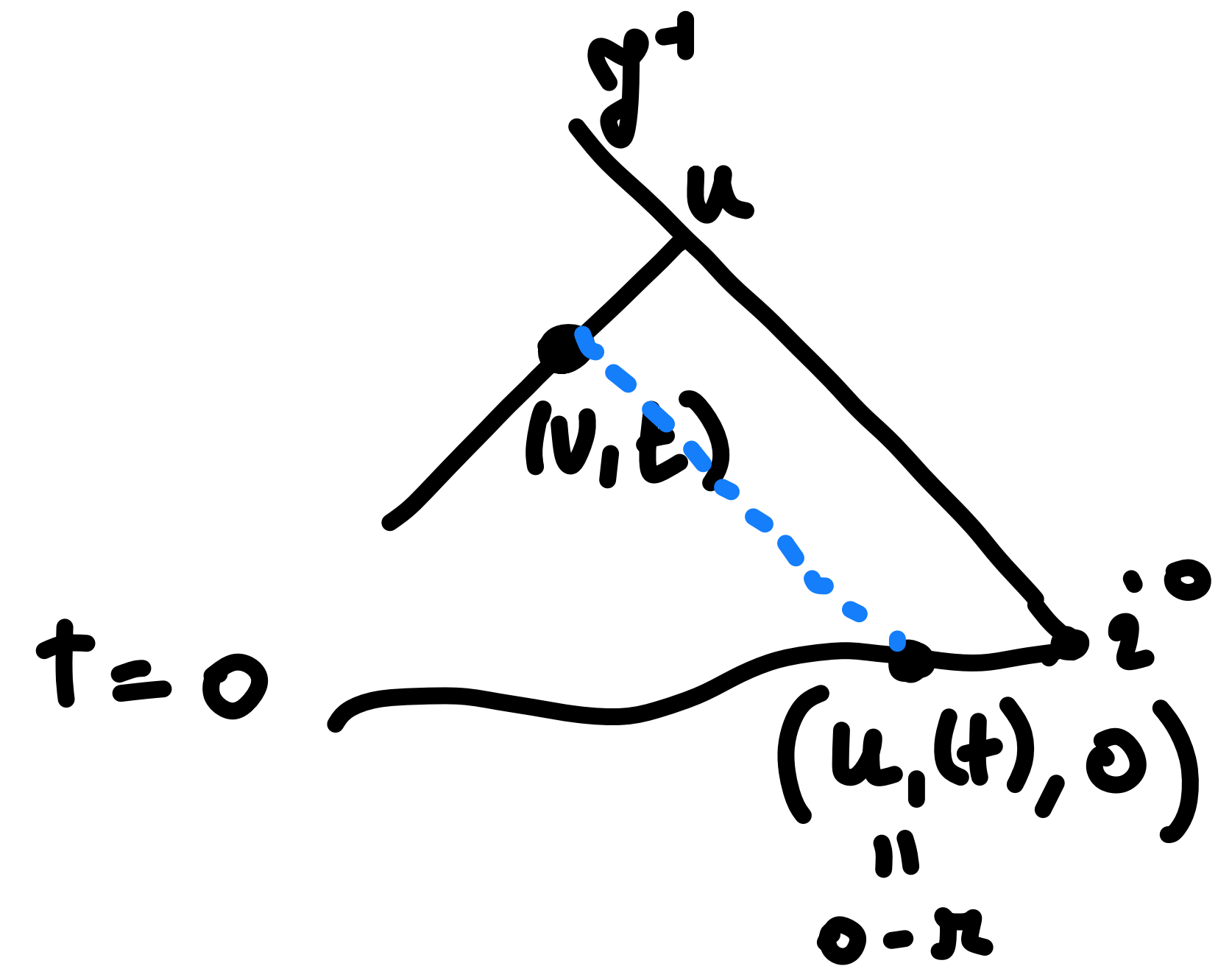
$$\sigma = \frac{C}{u} + \dots,$$

$$\Psi_3^0 = -\partial\partial_u\sigma, \Psi_4^0 = -\partial_u^2\sigma$$

$$\partial_u\Psi_2^0 = \partial\Psi_3^0 - \sigma\Psi_4^0 \Rightarrow \Psi_2^0 = -\frac{\partial^2 C}{u}$$

$$\partial_u\Psi_1^0 = \partial\Psi_2^0 - 2\sigma\Psi_3^0 \Rightarrow \partial_u\Psi_1^0 = -\frac{\partial^3 C}{u}$$

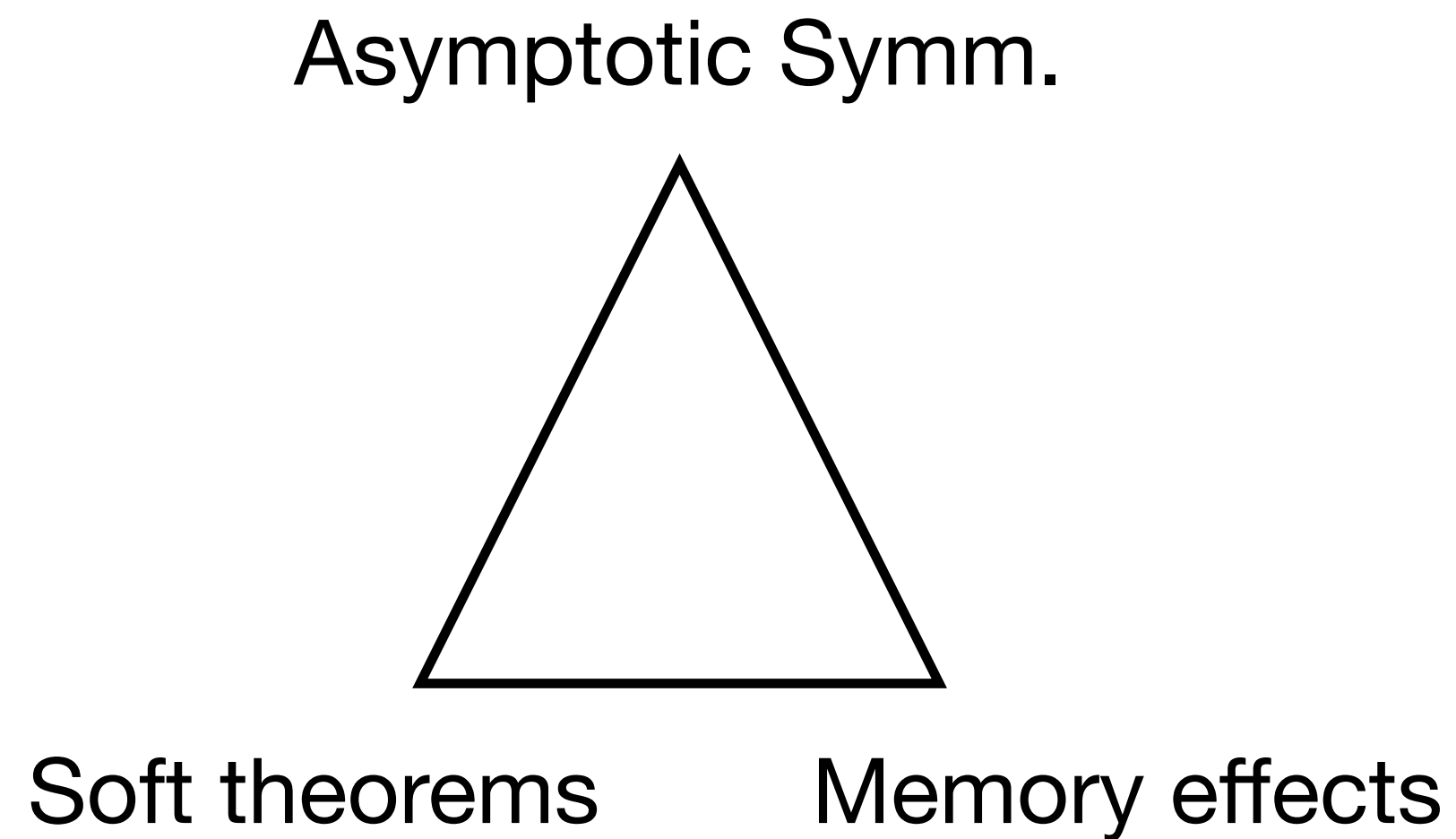
$$\partial := m_0^A \partial_A$$



$$u = t - r$$

$$\int_{u_1}^u \Psi_1^0 = -\partial^3 C (\ln u - \ln r)$$

# Infrared triangle



- Shown for leading infrared triangle (supertranslations, displacement memory, leading soft theorem)

[A. Strominger, '14; T. He, V. Lysov, P. Mitra and A. Strominger, '15; A. Strominger and A. Zhiboedov, '16]

- Use of charges/fluxes

- Log soft theorems and  $C_{AB} \sim \frac{1}{u}$  is shown to be related to superrotation Ward identity

[S. Agrawal, L. Donnay, K. Nguyen and R. Ruzziconi '23, S. Choi, A. Laddha and A. Puhm '24]

[L. Baulieu and T. Wetzstein '24]

**BUT** based on BC & symm. preserving the peeling ...

*What is the impact of violation of peeling on symmetries, charges, fluxes?*

# Symmetries in Log. Asympt. Flat

- BMSW is still a symmetries (the derivation we used didn't depend on the expansion of  $\gamma_{AB}$ )
- Charges (after renormalization), there is a contribution to supertranslation charges for generalized BMS

$$k_Y^{ur} = \bar{k}_Y^{ur} + 2\delta \left( \sqrt{q} Y^A D^B D_{AB} \right) - \frac{1}{2} D_C Y^C \sqrt{q} q^{AB} \delta D_{AB}$$

- Flux happens to be insensitive to the peeling!

# Equations of motion

# Evolution equations

## With peeling

- Flux balance law for the mass and the angular momentum

$$\partial_u \text{Re}(\Psi_2^0) = m_0^A \partial_A \Psi_3^0 - \sigma \Psi_4^0$$

-> Related to leading and sub soft theorem

- Tower of flux balance laws for  $E_{AB}^{n,0}(\Psi_0)$
- Related to the tower of (sub)^#- soft theorems and  $w_{1+\infty}$  algebra (in a sector)
- But not to local spacetimes symmetries (-> twistors)

# Evolution equations

## Log. Asympt. Flat

- Dressed eom for mass, angular momentum and  $E_{AB}^{n,0}$
- Conserved logarithmic branches

$$\partial_u D_{AB} = 0, \quad \partial_u E_{AB}^{n,n+1} = 0$$

- **NEW** flux balance law at sufficiently low order

$$\partial_u E_{AB}^{3,2} = \dots + \frac{3}{8} \partial_u C_{AB} C^{CD} E_{CD}^{1,2} - \frac{1}{8} E_{AB}^{1,2} \partial_u (C_{CD} C^{CD}) + \frac{3}{16} E_{C\langle A}^{1,2} \partial_u C^{CD} C_{B\rangle D}$$

*Sub log infrared triangle? (Wip)*

# Summary



# Summary

- Some physical phenomena require a violation of peeling (Log. Asympt. Flat BC)
- Symmetries of scattering also require a violation of peeling (Log. Asympt. Flat BC)
- BMS still preserve Log. Asympt. Flat BC
- Superrotation charge is modified (for generalized BMS)
- Fluxes are NOT modified -> the connection to (leading) logarithmic soft theorem can be made assuming peeling
- New tower of conserved quantities and flux balances law in the log-sector -> Connection to sub-log soft theorems?